Active Galactic Nuclei

- for most galaxies the luminosity is dominated by starlight
- small fraction (higher in the past) the nucleus of the galaxy is very bright

Active Galactic Nuclei (AGN)

Active Galactic Nuclei

Numerous different types of AGN:

- Seyfert galaxies
- Quasars
- Blazars
- Radio galaxies
- ..

Different observed properties: *all* thought to be powered by disk accretion on to a central supermassive black hole

First radio surveys of the sky done in 1950s

Many bright radio sources out of the plane of the Galaxy were found to coincide with star-like objects on photographic plates

Named "Quasi-stellar radio sources" (quasars)

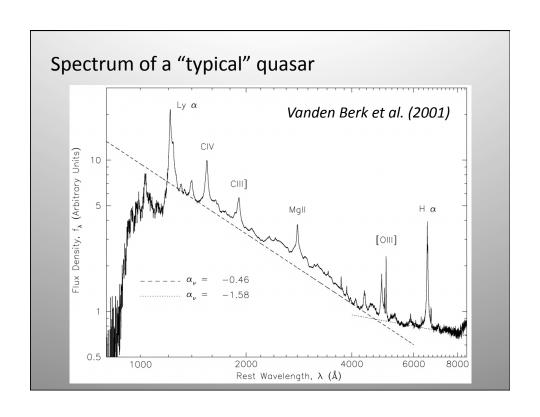
Nature initially a mystery

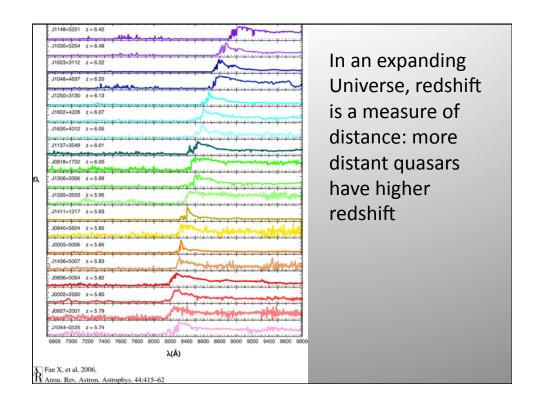
Quasars

Mystery was solved when it was realized that quasars were very distant objects, whose spectra were red shifted by the expansion of the Universe

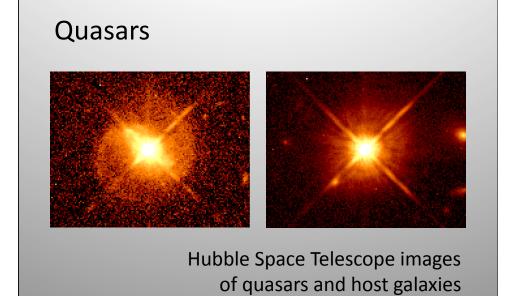
If a source emits light at wavelength λ_{emit} , and it is observed at λ_{obs} , define redshift z:

$$z = \frac{\lambda_{obs}}{\lambda_{emit}} - 1$$

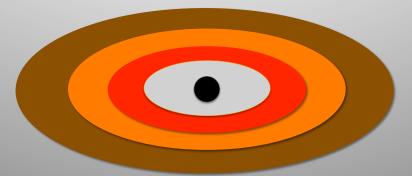




Large distance implies a very high luminosity: emission from the point source in the nucleus vastly outshines rest of the galaxy!



Interpret high luminosity as due to disk accretion of gas on to a supermassive black hole



Accretion rate \dot{M} - kg s⁻¹ (Solar masses per year)

Quasars

Conversion of rest mass of accreting gas to radiated energy is about 10% for disk accretion on to a black hole, i.e. for 1 kg accreted get:

$$E = 0.1 \text{ kg} \times c^2$$
 of energy in the form of radiation

Generally, luminosity:

$$L = \varepsilon \dot{M}c^2 \approx 0.1 \dot{M}c^2$$
radiative efficiency of accretion

Example: suppose a billion Solar mass black hole is accreting gas at 1 Solar mass per year

$$\dot{M} = \frac{2 \times 10^{30} \text{ kg}}{3.16 \times 10^7 \text{ s}} = 6.3 \times 10^{22} \text{ kg s}^{-1}$$

Expect a luminosity of:

$$L \approx 0.1 \dot{M}c^2 \approx 5.7 \times 10^{38} \text{ Watts}$$

Quasars

$$L \approx 0.1 \dot{M}c^2 \approx 5.7 \times 10^{38} \text{ Watts}$$

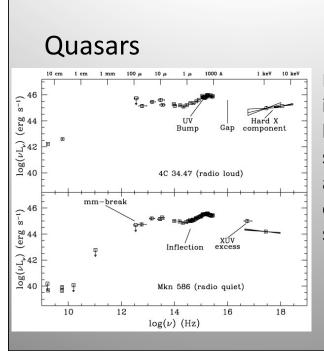
Compare to luminosity of Sun, 3.8 x 10²⁶ Watts

Luminosity from the accreting black hole is equal to 1.5 trillion (1.5 x 10^{12}) stars like the Sun!

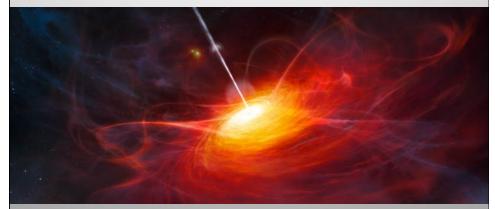
Consequence of the high efficiency of black hole accretion vs nuclear fusion (10% vs 0.7%)



Hard even to see the stars in the galaxy: the nucleus is so bright

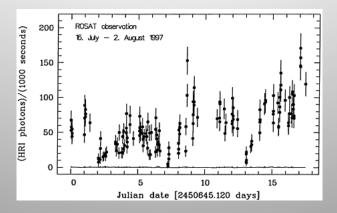


Emit most energy in UV and optical light, but also shine brightly across the whole electromagnetic spectrum



Luminosity from hot gas in the accretion disk, also from a *jet* possibly powered by black hole spin energy

Maximum size from variability



Quasars (and other AGN, this is a Seyfert) are seen to vary on time scales of a day or less

Maximum size from variability



For a source of size one light day, fastest time for a signal to cross the source is one day



Don't expect to see co-ordinated variations on less that 1 day time scale

Maximum size from variability

Observe variability on 1 day time scale

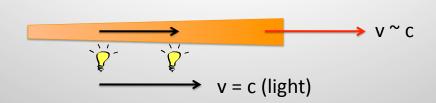
Infer quasars are less than 1 light day across

$$L \sim (24 \times 3600 \text{ s}) \times 3 \times 10^8 \text{ ms}^{-1}$$

$$= 2.6 \times 10^{13} \text{ m} = 170 \text{ AU}$$



luminosity of a whole galaxy originates from region ~size of the Solar System!



A loophole: if the emitting source is moving toward us at close to the speed of light:

- source *almost* keeps pace with light emitted earlier
- "slow" variability appears faster to us
- source can be bigger than light travel time argument suggests

A mystery:

- at z = 7, time since the Big Bang was less than a billion years
- very few galaxies had formed by that time
- how did a billion Solar mass black hole grow that large in the time available?