

MOLECULAR SPECTRA

Pure rotational spectra (no elec, vib transition)

Selection rules for em/abs of elec dipole rad

1. Molecule must have permanent dipole moment

$$\vec{d} = \vec{d}_{\text{nuc}} + \vec{d}_{\text{elec}}$$

2. $\Delta J = -1$ em

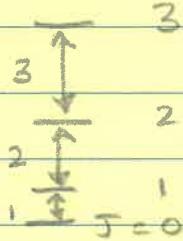
+1 abs

3. $\Delta J = 0$ also allowed

for asymmetric tops ($I_A < I_B < I_C$)

Note that $\vec{J} = \vec{K} + \vec{S}_{\text{elec}}$

So $\Delta K = \Delta J$ in pure rot transition.



$\Delta E_{01} = \Delta E_{12} = \Delta E_{23}$

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Rotational spectra of homonuclear molecules

H_2, D_2, O_2, \dots in ground elec term $\Lambda = 0$

have no dipole moment.

But do radiate by quadrupole radiation

$\Delta K = \Delta J = -2$ em

+2 abs

Ro-vibrational spectra

= vibrational transition $\nu_u - \nu_L$

together with rotational transition.

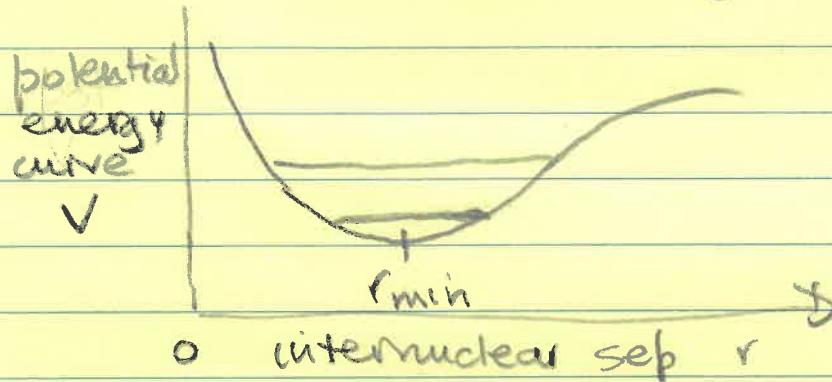
Polarizability eg. oscillating em photon field

Apply external elec field \vec{E} to molecule.

Induced dipole moment is Herzberg p. 82

$$\vec{d} = \alpha \vec{E}$$

α polarizability (a matrix)



Expand α as Taylor series about $r = r_{\min}$:

$$\alpha = \alpha_0 + \alpha_1(r - r_{\min}) + \dots$$

Then

$$\langle \psi_u | \vec{d} | \psi_L \rangle = \vec{E} \left\{ \alpha_0 \langle \psi_u | \psi_L \rangle + \alpha_1 \langle \psi_u | r - r_{\min} | \psi_L \rangle + \dots \right\} \vec{E}$$

↑ ↑
quantum STO
wavefunctions

Linear term $\neq 0$ only for $\nu_u = \nu_L + 1$.

So to extent that

(i) higher terms of polarizability small;

(ii) STO good approximation,
get vibrational selection rule

$$\Delta\nu = -1 \text{ em}$$

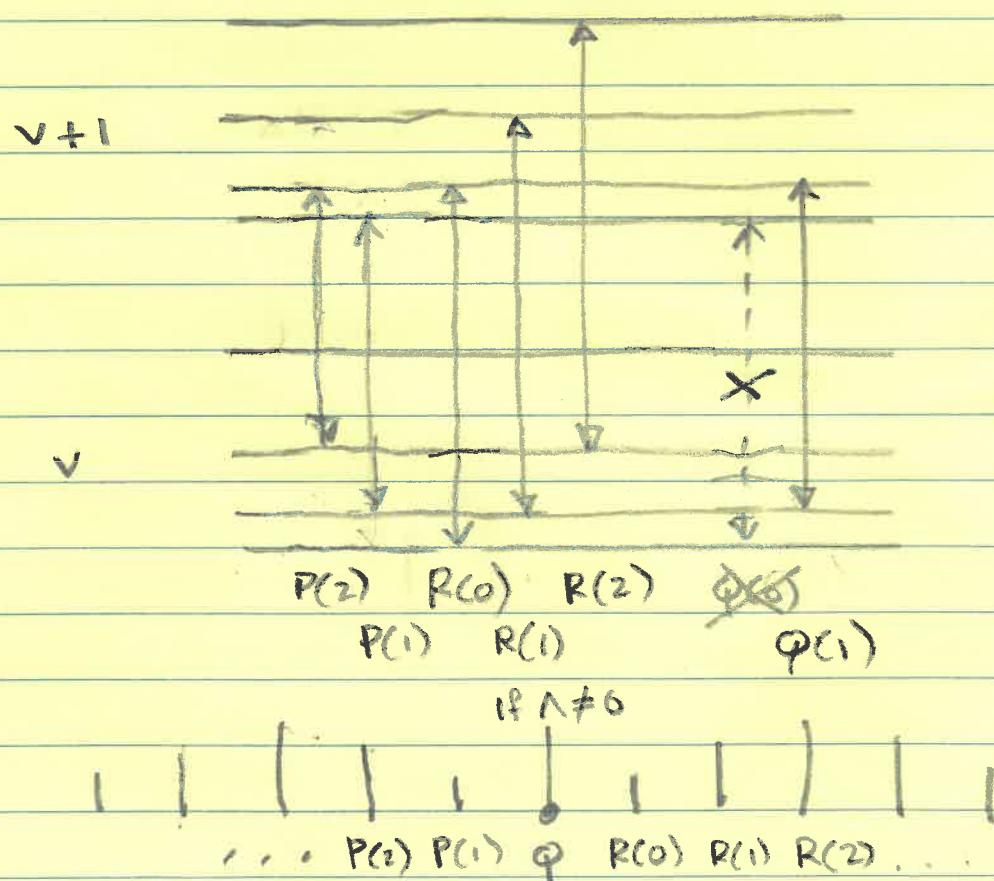
$$+1 \text{ abs}$$

Selection rules for em/abs of elec dipole rad
for ro-vibrational spectra:

1. Linear polarizability $\alpha \neq 0$ } vib rules
2. $\Delta v \approx \pm 1$
3. $\vec{d}_{\text{perm}} \neq 0$
4. $\Delta J = \pm 1 \quad (\Lambda = 0)$
5. $\Delta J = 0, \pm 1 \quad (\Lambda \neq 0) \text{ but } J=0 \neq 0$ } rot rules

Notation for transition:

$$(v_u - v_L) \left\{ \begin{array}{ll} P(J_L) & J_u = J_L - 1 \\ Q(J_L) & J_u = J_L \quad (\Lambda \neq 0) \\ R(J_L) & J_u = J_L + 1 \end{array} \right.$$



frequency →

arXiv 1509.06607
Fig 2

Transitions with $\Delta J = \pm 2$

occur in:

- homonuclear molecules e.g. H_2 , N_2 , O_2
which have no permanent elec dipole moment
- "Raman" scattering
= scattering of photons of any frequency



Nomenclature

$$(v_u - v_L) \left\{ \begin{array}{ll} O(J_L) & J_u = J_L - 2 \\ P(J_L) & J_u = J_L - 1 \\ Q(J_L) & J_u = J_L \quad J_L \neq 0 \\ R(J_L) & J_u = J_L + 1 \\ S(J_L) & J_u = J_L + 2 \end{array} \right.$$

1509.06607 1507.06650 Fig 1

12.5

Example ro-vibrational lines

arXiv: 1304.4961

Class: (based on slope of SED)

^{12}CO 2-1

1. IR richest \downarrow evolve

^{13}CO 1-0

2. \downarrow evolve

^{12}CO 1-0

"CO ice" means? at $4.67\mu\text{m}$

Proto planetary disks - ESO VLT

arXiv: 1211.6750 NGC 1275 = Per A,

$$\text{H}_2(1-0) \text{ S} \left(\begin{array}{c} 0 \\ 1 \\ 2 \\ 3 \\ 4 \end{array} \right) \quad v=1, J=\frac{2}{3}, \frac{1}{4} \rightarrow v=0, J=\frac{1}{2}$$

$$(2-1) \text{ S} \left(\begin{array}{c} 0 \\ 1 \\ 2 \\ 3 \\ 4 \end{array} \right)$$

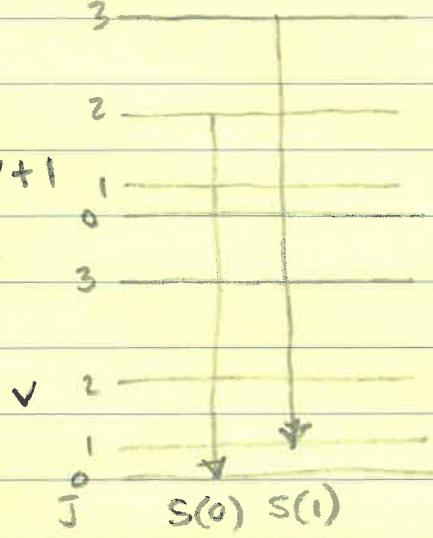
$$(3-2) \text{ S} \left(\begin{array}{c} 0 \\ 1 \\ 2 \\ 3 \\ 4 \end{array} \right)$$

H I Brγ $n=7 \rightarrow n=4$

Brackett

He I $2p^1P_1 \rightarrow 2s^1S_0$

$\text{H}_2(1-0) \text{ S}(0)$ $2.122 \mu\text{m}$



(12.6)

Electronic / vib / rot transitions

Molecule is approx stationary

while electron undergoes transition



Franck-Condon principle for vib transition

Hönl-London factor " rot "

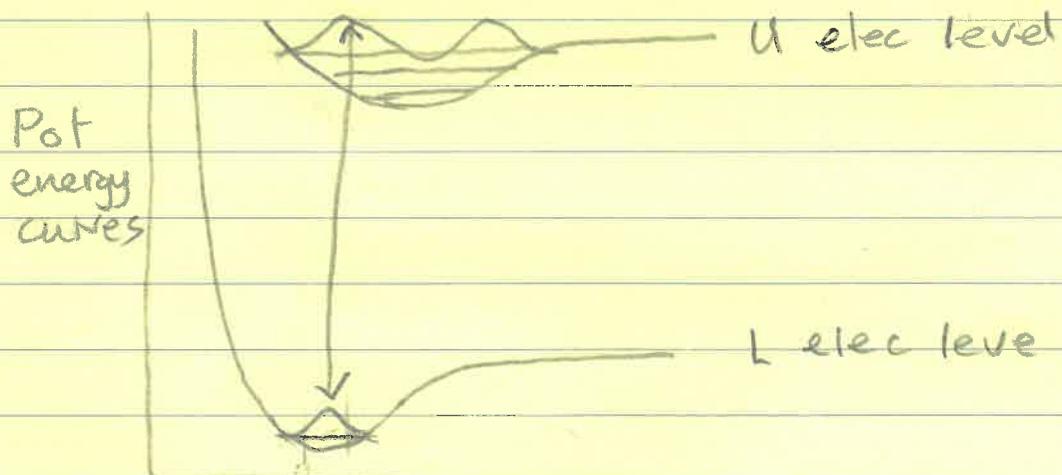
Franck-Condon principle

In elec/vib/rot transition,

vibrational probability is

$$\text{Prob} (v \leftrightarrow v') = |\langle v' | v \rangle|^2 \\ = |\text{overlap of vib w.f.s}|^2$$

Note $\sum |\langle v' | v \rangle|^2 = 1$



(12.7)

Hönl - London factors

In elec/vib/rot transition,
rotational probability is

$$\text{Prob } (J', K', S, \Lambda' \leftrightarrow J, K, S, \Lambda) = f(J' K' S' \Lambda' | J K S \Lambda)$$

$S' = S$ Hönl - London factor

Note factors have stat. weight built in,
so factor is same for em & abs.
See notes honl.ps.

Diatomic molecules: elec dipole selec rules

Herzberg p.240.

1. $\Delta \Lambda = 0, \pm 1$
 2. $\Delta J = 0, \pm 1$
- but $J = 0 \leftrightarrow 0$

} emitted photon
has unit ang mom

3. Total parity (elec & nuc) changes
because elec dipole phot has odd parity.

4. $\Delta S = 0$ because in $\langle n' \Lambda' | \vec{d} | n \Lambda \rangle$
elec spin dipole op \vec{d} indept of spin.

5. If $\Lambda = 0 \leftrightarrow 0$, then $\Sigma^+ \leftrightarrow \Sigma^+$ or $\Sigma^- \leftrightarrow \Sigma^-$

6. If same charge nuclei, then $\Lambda'_g \leftrightarrow \Lambda_u$

In addition,

Hund case (a)

7. $\Delta \Sigma = 0$ because elec spin unchanged
proj of elec spin along axis

$$\text{Hence } \Delta \underline{\Sigma} = \Delta \Lambda = 0, \pm 1 \\ \equiv \Lambda + \Sigma$$

8. $\Delta J \neq 0$ if $\Sigma = 0 \leftrightarrow 0$

Hund case (b)

9. $\Delta K = 0, \pm 1$

but $\Delta K \neq 0$ if $\Lambda = 0 \leftrightarrow 0$

Main branch has $\Delta J = \Delta K$ ($\vec{J} = \vec{F} + \vec{S}_{\text{elec}}$)

satellite branches $\Delta J \neq \Delta K$

Total (elec & nuc) parity = $\pm (-)^K$

For $\Lambda = 0$, \pm as Σ^\pm

For $\Lambda \neq 0$, each level is Λ -doubled, with even/odd parities, eg. Π^\pm .

Thus allowed transition are

$$+ \leftrightarrow + \quad \Delta K = \pm 1$$

eg $\Sigma^+ \leftrightarrow \Sigma^+$, $\Pi^+ \leftrightarrow \Pi^+$, $\Sigma^+ \leftrightarrow \Pi^+$

$$+ \leftrightarrow - \quad \Delta K = 0$$

eg $\Pi^+ \leftrightarrow \Pi^-$, $\Sigma^+ \leftrightarrow \Sigma^-$.

Example: elec/vib/rot transitions of H₂.

Allowed 1s - 2p transitions

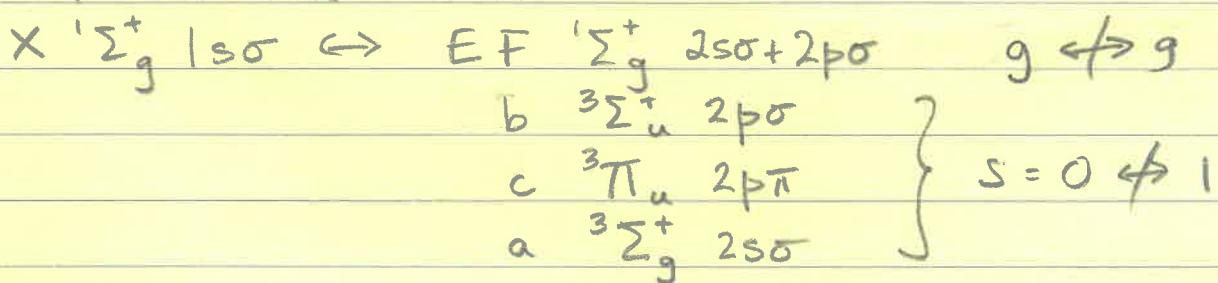
"Lyman" band $\times ^1\Sigma_g^+ 1s\sigma \leftrightarrow B ^1\Sigma_u^+ 2p\sigma$

"Werner" band $\leftrightarrow C ^3\Pi_u^+ 2p\pi$

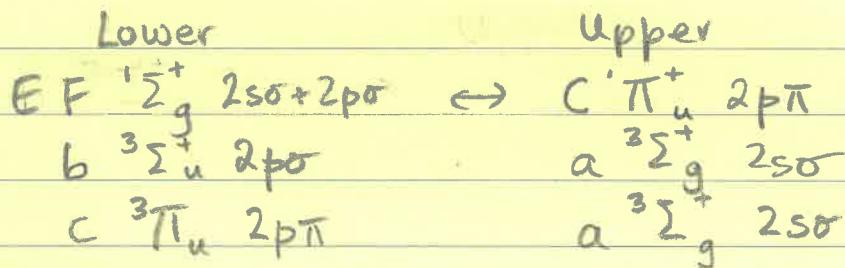
arXiv: 0804.4761 Fig 7 p 12 - astro-ph/0507581 Fig 3. p 30

Abs line from $v=0$ and $J=0, 1, 2$.

Forbidden transitions:



Allowed 2 - 2 transitions



Note

$b ^3\Sigma_u^+ 2p\sigma \leftrightarrow c ^3\Pi_u^+ 2p\pi$ forbidden $u \leftrightarrow u$.
 But lowest $c ^3\Pi_u^+$ state has nowhere else to decay to. It is "metastable" and decays to $b ^3\Sigma_u^+$.

Here $\Delta S = 0$ rule beats $g \leftrightarrow u$ rule

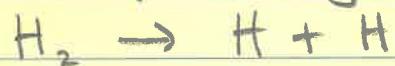
Notation

$$\text{Lyman } L(v_u - v_L) \left\{ \begin{array}{l} P(J_L) \\ R(J_L) \end{array} \right.$$

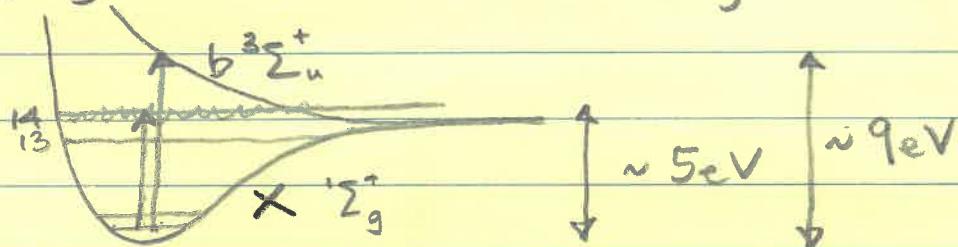
$$\text{Werner } W(v_u - v_L) \left\{ \begin{array}{l} \\ \end{array} \right.$$

Dissociation of H_2 binding energy 4.52 eV.

Various pathways



(1) Collisional exc of $v = 14$ vib level of ground elec term $X^1\Sigma_g^+$ unbound



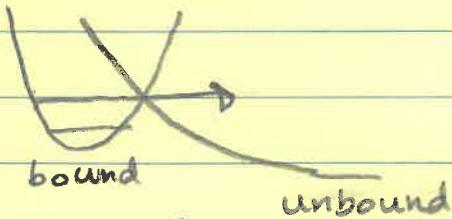
Unbound levels are broad $\Delta E \Delta t \sim \hbar$ with $\Delta t \sim 1$ molecular vib time.

(2) Collisional exc of $b^3\Sigma_u^+$ triplet, which is unbound.

(3) UV photons excite singlet levels (eg B, C) which decay radiatively to unbound levels, eg $X^1\Sigma_g^+ v \geq 14$.

Predissociation

is transfer from bound elec/vib state to unbound elec/vib state



Association of H₂

Radiative association



is dipole forbidden, very slow.
So how does H₂ form?

(1a) H + H → H₂ on surface of dust grain

(1b) Three-body recombination



Both these are reactions at "high" density.

(2) At T \gtrsim 1500 K, where H is partially ionized,



Why require hi T for H⁻? Need e for H + e \rightarrow H⁺

Primordial H₂

$n \gtrsim 10^4 \text{ cm}^{-3}$ for TE pop of rot levels.
HD provides additional coolant.



$$n \gtrsim 10^8 \text{ cm}^{-3}$$



Abel et al 1997

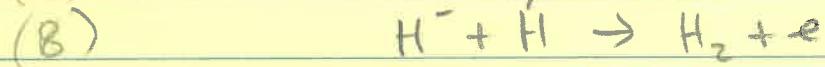
New Astron 2 181

Galli & Palla 98

AA 335 403

Glover & Abel 2008

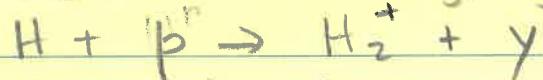
MN 388 1627



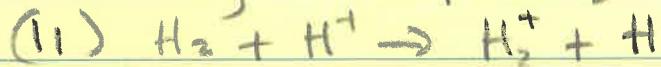
is dominant channel.

Alternative is

H₂⁺ is more tightly bound than H⁻



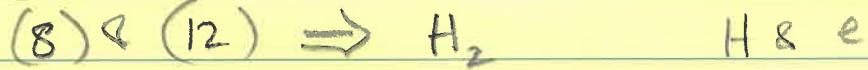
Cooling is H + H₂ exciting rovib.



(12.14)

Minimal model

- (1) $H + e \rightarrow H^+ + 2e$ coll ion
- (2) $H^+ + e \rightarrow H + \gamma$ rad rec
- * (8) $H + H^- \rightarrow H_2 + e$ H_2 form
- (11) $H_2 + H^+ \rightarrow H_2^+ + H$
- * (12) $H_2 + e \rightarrow 2H + e$
- * (7) $H + e \rightarrow H^- + \gamma$
- (16) $H^- + H \rightarrow 2H + e$
- (14) $H^- + e \rightarrow H + 2e$



All H^- goes to H_2 via (8)

H^- created via (7)

Does not affect e abundance - e is catalyst.